

Luminescence chronology of cave sediments at the Atapuerca paleoanthropological site, Spain

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A B S T R A C T

Ascertaining the timing of the peopling of Europe, after the first out-of-Africa demographic expansion at the end of the Pliocene, is of great interest to paleoanthropologists. One of the earliest direct evidences for fossil hominins in western Europe comes from an infilled karstic cave site called Gran Dolina at Atapuerca, in a stratum ~1.5 m below the Brunhes–Matuyama (B–M) geomagnetic boundary (780 ka) within lithostratigraphic unit TD6. However, most of the meters of fossil- and tool-bearing strata at Gran Dolina have been difficult to date. Therefore, we applied both thermoluminescence (TL) and infrared-stimulated-luminescence (IRSL) multi-aliquot dating methods to fine-silt fractions from sediment samples within Gran Dolina and the nearby Galería cave site. We also applied these methods to samples from the present-day surface soils on the surrounding limestone hill slopes to test the luminescence-clock-zeroing-by-daylight assumption. Within the uppermost 4 m of the cave deposits at Gran Dolina, TL and paired TL and IRSL ages range stratigraphically from 198 ± 19 ka to 244 ± 26 ka. Throughout Gran Dolina, all luminescence results are stratigraphically self-consistent and, excepting results from two stratigraphic units, are consistent with prior ESR-U-series ages from progressively deeper strata. Thermoluminescence ages culminate at 960 ± 120 ka approximately 1 m below the 780 ka B–M boundary. At Galería, with one exception, TL and IRSL ages range stratigraphically downward from 185 ± 26 ka to 503 ± 95 ka at the base of the lowermost surface-inwash facies. These results indicate that TL and (sometimes) IRSL are useful dating tools for karstic inwash sediments older than ca. 100 ka, and that a more accurate chronostratigraphic correlation is now possible among the main Atapuerca sites (Gran Dolina, Galería, Sima de los Huesos). Furthermore, the oldest TL age of ca. 960 ka from Gran Dolina, consistent with biostratigraphic and paleomagnetic evidence, implies a probable numeric age of 900–950 ka for the oldest hominin remains (~0.8 m below the TL sample). This age window suggests a correspondence to Marine Isotope Stage (MIS) 25, a relatively warm and humid interglaciation.

Keywords:

Atapuerca

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Introduction

The Atapuerca paleoanthropological sites, near Burgos in northern Spain, contain the world's most extraordinary accumulation of Pleistocene fossil humans and related stone artifacts (e.g., Arsuaga et al., 1993, 1999a,b; Carbonell et al., 1995, 1999; Bermúdez de Castro et al., 1997, 1999, 2004; Parés et al., 2006). In particular, the infilled karstic cave site called Gran Dolina (TD) has yielded the heretofore oldest direct evidence of hominins in Europe west of the

Republic of Georgia. More than a hundred hominin remains have been recovered so far from the so-called Aurora stratum within lithostratigraphic unit TD6, ~1.5 m below the Brunhes–Matuyama (B–M) geomagnetic boundary (Parés and Pérez-González, 1999).

These remarkable finds have generated discussion and controversy about the origin of modern Europeans (e.g., Delson, 1997; Ambrose, 2001; Manzi, 2004; Parfitt et al., 2005; Roebroeks, 2005; Dennell and Roebroeks, 2005; Carbonell and Rodríguez, 2006). Speculation about the possible migration routes of human ancestors into Europe (e.g., Aguirre and Carbonell, 2001; Balter, 2001; Dennell, 2003) has recently been reinterpreted by Dennell and Roebroeks (2005), who argued that the extant “out-of-Africa-1” (e.g., Tattersall, 1997) hypothesis should be replaced by a scenario in

which early hominins occupied Asiatic grasslands by 3.0–3.5 Ma, followed by “two-way traffic,” which would be responsible for the presence of the genus *Homo* across central-southern Asia by ca. 1.8 Ma. In this context, migratory origins (Ambrose, 2001; Balter, 2001; Cann, 2001; Culotta et al., 2001; Gibbons, 2001a,b; Marshall, 2001a,b; Pennisi, 2001; Stumpf and Goldstein, 2001) of the fossil humans in Spain, now known to have been present at least as early as the time window 0.78–1.80 Ma (Parés et al., 2006; Scott et al., 2007), take on new significance. For example, new questions can be considered. If hominins were present in Spain as early as 1.8 Ma, what was their relationship, if any, to hominins apparently present at about that time across central-southern Asia? If hominins were present at Gran Dolina close to 1 Ma, as our data imply, what was their relationship, if any, to those hominins present in the same area (Parés et al., 2006) at earlier times (e.g., 1.2–1.8 Ma)?

The most important sites at Atapuerca are Gran Dolina (site TD) and Galería (TG) in an abandoned railway cut or trench (La Trinchera), as well as Sima de los Huesos (SH; “Pit of the Bones”) approximately 400 m away (Fig. 1). English-language histories of the exploration and discovery of these remarkable sites have been presented by Aguirre et al. (1990), Arsuaga et al. (1997a,b), Balter (2001), and Carbonell et al. (1999), with popular accounts given by Arsuaga (2003) and Kunzig (1997). The increasing importance of a third site within the railway trench, Sima del Elefante (site TE; Fig. 1), is recognized (Rosas et al., 2004; Parés et al., 2006). There, paleomagnetic and biostratigraphic results have placed the age of lithic tools (hence the presence of hominins) in the early Pleistocene (0.78–1.77 Ma).

Dating the remains from La Trinchera and their enclosing sediments has been unusually difficult. The main reason is that the sediments are infill deposits in Pleistocene caverns within a karstic terrain of the Sierra de Atapuerca, and thus contain no heated minerals amenable to Ar–Ar and fission-track dating. However, dating authigenic minerals by Ar–Ar methods (e.g., Dickin, 1997; McDougall and Harrison, 1999; Hagen et al., 2001) may be possible, and dating unheated quartz grains by cosmogenic-nuclide methods (e.g., Dickin, 1997; Gosse and Phillips, 2001; Anthony and Granger, 2007) is possible (Carbonell et al., 2008). Several of the deposits are within the 50–400-kyr age range of conventional U–Th-series dating (e.g., Dickin, 1997; Ku, 2000), and within the somewhat longer age range of electron-spin-resonance (ESR) dating of tooth enamel (e.g., Rink, 1997). The relative-dating tool of geomagnetic-reversal stratigraphy (e.g., Verosub, 2000) has been applied at Atapuerca. Thus, the most reliable and critical datum for assigning an age to the oldest hominin remains at Gran Dolina has been the determination of the B–M geomagnetic-reversal boundary (age 781 ± 3 ka; Horng et al., 2002) in upper stratigraphic unit TD7 (Parés and Pérez-González, 1999), ~1.5 m above the hominin-bearing horizon in unit TD6.

Because of the importance of these sites and of the limited application of other dating techniques, we have applied those multi-aliquot luminescence sediment-dating methods (e.g., Aitken, 1985, 1998) that were available to us at the time of commencement of this project (1998). We thought that the more recently developed methods of single-aliquot and single-grain photon-stimulated-luminescence dating (Aitken, 1998) (also termed optical or photonic dating) of quartz (e.g., Murray and Olley, 2002; Berger et al., 2003; Bøtter-Jensen et al., 2003) were unlikely to be useful here because all our stratigraphically important samples were thought to be greater than ca. 100 kyr in age. For sediments having typical concentrations of U, Th, and K, precise single-aliquot-quartz dating is usually limited to the last ca. 150 kyr (Murray and Olley, 2002). However, Rhodes et al. (2006) reported that one calcareous sandy-silt sample somewhat below the B–M geomagnetic-reversal boundary, at an archeological site near Casablanca in Morocco, yielded an optical age estimate of 990 ± 210 ka, while some

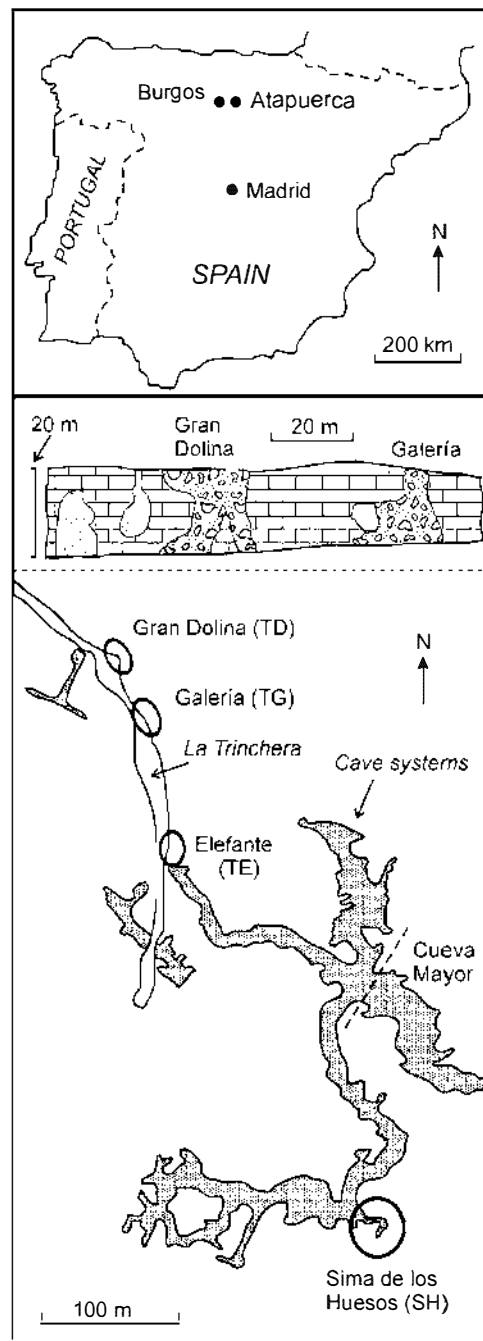


Fig. 1. Location of the Atapuerca sites (top), schematic profile of the Gran Dolina and Galería cave deposits (center), and a plan view of the sites and cave system (lower). The profile view shows mixed sediments infilling passages within the limestone bedrock, with an unfilled cavity at the left side of Galería. In the plan view, the subsurface caves are shaded.

samples stratigraphically above the B–M boundary gave optical ages on quartz of ca. 400–500 ka. Importantly, the dose rates in all their sediments were unusually low, 0.4–0.8 Gy/ka, thus permitting quartz dating well beyond ca. 150 ka. The recent demonstration of single-aliquot-quartz optical ages up to ca. 500 ka for a karstic cave setting in Australia (Prideaux et al., 2007a) similarly depends on low dose rates in the enclosing sediments (~0.6 Gy/ka).

Gran Dolina and Galería are the main subjects of this paper because, of the three sites available to us in the late 1990s (TD, TG, SH), only TD and TG have sediments that might meet the main requirement of luminescence sediment dating that mineral grains be exposed to daylight not too long before burial.

Geological setting and samples

The geological setting of the endokarst and the nature of the deposits have been described by Carbonell et al. (1999), Parés and Pérez-González (1999), and Pérez-González et al. (1999, 2001). From the viewpoint of luminescence dating, there are three important traits in the setting and the deposits: (1) the bedrock is Mesozoic limestone, and the karstic structure was developed no later than the early Pleistocene (e.g., Pérez-González et al., 1999, 2001; Parés et al., 2006); (2) upslope of the filled TD and TG cave entrances, the bedrock surface is covered by patches of thin (10–50 cm thick) terra-rossa soil (e.g., Jenny, 1980; Martini and Chesworth, 1992); (3) the cave passages are filled with both a reddish matrix (exterior or “entrada” facies deposits) and an underlying (basal) tan-beige matrix (“interior” facies) (Pérez-González et al., 1999, 2001).

The 18-m-thick exposure of interior and exterior facies deposits at TD is divided into 11 stratigraphic levels (Fig. 2, left): TD1–TD11,

bottom to top (e.g., Parés and Pérez-González, 1995, 1999). The ~16-m exposure at TG is divided into six stratigraphic levels (Fig. 2, right) (G1–G6, bottom to top; Pérez-González et al., 1995, 1999, 2001), with G1 being an interior facies deposit, and G2 and G3 containing fossils and evidence of hominin occupation.

A limestone character of the bedrock ensures that any fragments of bedrock that are dissolved or finely comminuted will contribute a very small or negligible proportion of feldspars and quartz to the sediment matrix. Quartz and feldspar grains carry most of the luminescence signal in typical sediments. Our objective was to date sediment greater than 100 ka in age. Although the exterior facies deposits have ~12–50% quartz (Pérez-González et al., 1995, 1999), and although quartz provides a highly stable luminescence signal (e.g., Aitken, 1985, 1998), the upper age limit of quartz in normal sediments is ca. 200 ka or less. Alternatively, feldspar concentration in these cave deposits was estimated to be only ~5% (Pérez-González et al., 1995, 1999). However, because feldspar thermoluminescence (TL) is ~20–50

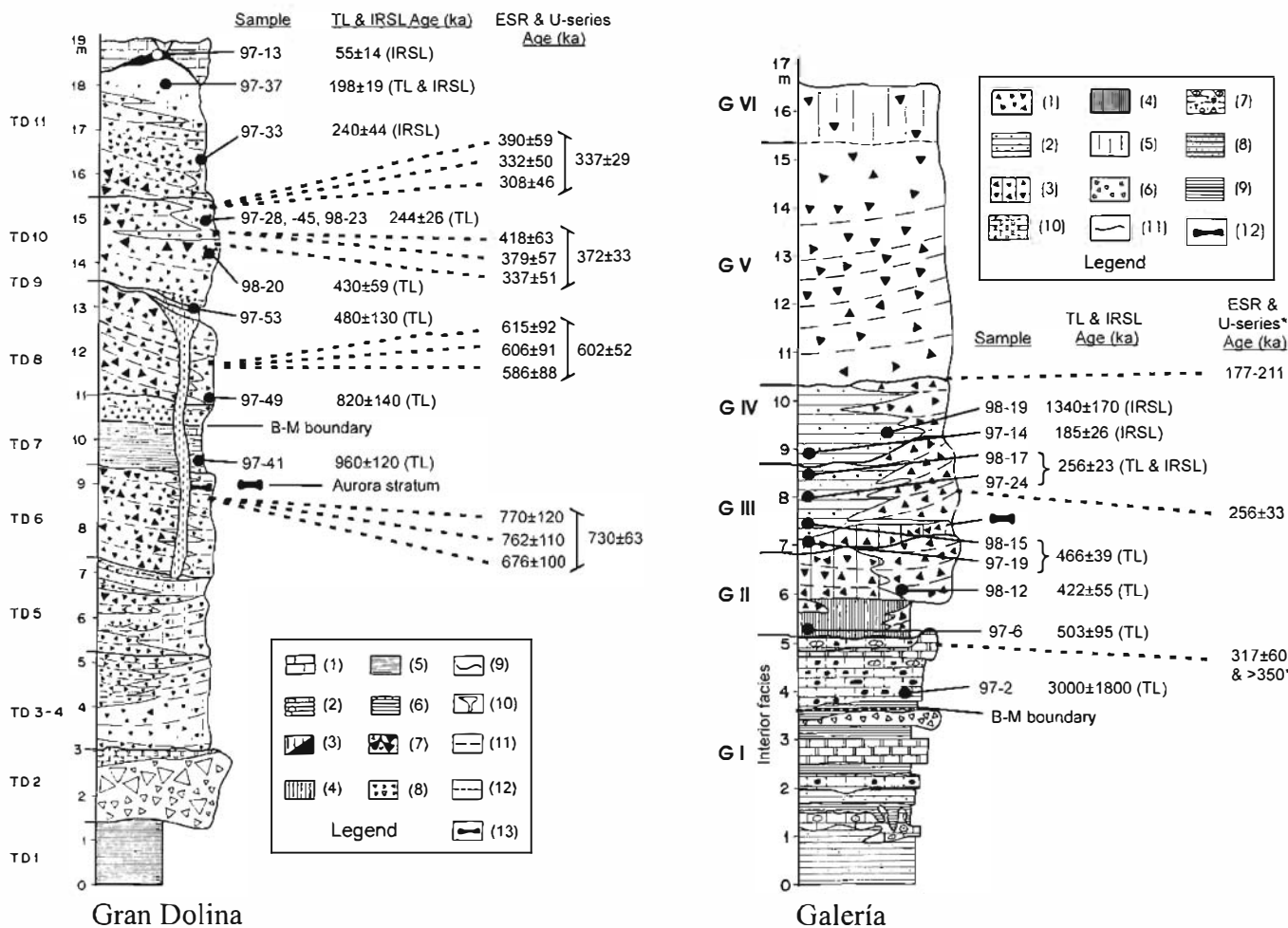


Fig. 2. Composite stratigraphic profiles of the cave deposits at (left) Gran Dolina (updated from Parés and Pérez-González, 1999) and (right) Galería, with the locations of the luminescence samples, luminescence ages (double asterisks [**] denote TL and IRSL ages from Table 2), and independent ages. For Gran Dolina, the analytical errors for the three oldest ESR-U-series ages are here rounded from the reported (Faluères et al., 1999) three to two statistically significant digits, as is more appropriate for such data (e.g., Topping, 1962). A weighted mean and standard error of mean (Topping, 1962) are shown for each ESR-U-series data group. For Gran Dolina, the legend is: (1) Mesozoic limestone at the Dolina roof; (2) speleothem; (3) lutites, clay loam/terra rossa; (4) guano of bats; (5) laminated loamy clays; (6) calcilutites and calcarenites; (7) gravels and boulders, clastic flow; (8) deposition of fallen boulders; (9) main stratigraphical discontinuity; (10) secondary unconformity and loamy-clayey-sandy filling; (11) Brunhes–Matuyama boundary; (12) disappearance of *Mimomys savini* and first occurrence of *Iberomys breccienensis*; (13) location of the Aurora stratum (containing the oldest hominin remains) at the top of TD6. For Galería, the legend is: (1) gravels and boulders; (2) alternation of small gravels and clay loam; (3) gravels and clay loam; (4) guano and clay loam; (5) soil; (6) small gravels; (7) sandy clay loam with clay clasts or carbonates; (8) very fine sands and loams with laminar structure; (9) sandy clays with laminar structures; (10) speleothem; (11) main stratigraphical discontinuity; (12) a fragment of a hominin cranium (right parietal) was recovered from the base of GIII (G3 hereafter) (Arsuaga et al., 1999c), but it is not well preserved and no species assignment has been made.

times brighter than an equivalent amount of quartz (Wintle, 1982; Berger, 1984; Aitken, 1985), we expected TL signals from poly-mineral, fine-silt-sized samples to be dominated by feldspar signals. These cave deposits are also suitable for multi-aliquot TL or infrared-stimulated-luminescence (IRSL) dating because there is little likelihood of significant contamination by light-shielded quartz or feldspar grains that could be released into the matrix by dissolution of local limestone, either in situ or during laboratory preparation of samples containing small clasts of limestone. The concentration of feldspar and quartz within the local limestone is a factor of at least 40–500 less than in the cave sediments (Aleixandre and Pérez-González, 1999).

The strikingly reddish coloration of the karstic infill matrix overlying the lowest cave units (which are tan-colored) attests to the origin of the infill matrix as downslope wash from the past terra-rossa surface soil (Pérez-González et al., 1995, 1999). In this context, the preservation of a remnant terra rossa on the upslope bedrock surface provides a modern analogue to the likely past upslope material that periodically washed downslope into the cave entrances, mixing with small and large limestone breccia, to gradually fill the TD and TG caves.

Most of the upslope sediment that washed into the caves was likely transported during energetic events. Because mass transport was the main mechanism in the trinchera (Pérez-González et al., 1999), it is unlikely that many or any of the fine-silt-sized mineral grains would have been exposed to sufficient daylight to significantly reset the luminescence clock during transport. In general depositional settings, floodwaters and episodic transport of slurries typically shield most grains from light. There are many published examples illustrating gross age overestimation when such alluvial sediments are dated by multi-aliquot TL and IRSL. Therefore, before we could fully interpret TL and IRSL ages from the matrix of the cave deposits, we needed to establish that upslope sediment likely had small residual luminescence ages before being transported into the caves. For this reason we collected samples from the present thin sediments on the limestone bedrock upslope of the caves.

Our hypothesis was that the reddish matrix of the exterior facies deposits in TD and TG contains a significant fraction of eolian silt grains. Continually added to surface sediments (see online Supplementary material), such grains would carry a small relict luminescence age. Because the youngest deposits in the caves were thought to exceed ca. 100 ka, and since a typical analytical uncertainty in luminescence age estimates is ~10%, we require any relict or inherited age from the inwash sediment to be less than 10 kyr (but could be larger for older samples) for samples older than 100 ka.

Within the various cave deposits, we collected blocks of matrix during daylight, and opaque bags of friable matrix sample during diurnal darkness, at stratigraphic locations shown in Fig. 2. Sample collection occurred in the summers of 1997 and 1998. All outer portions (~5–10 mm) of daylight-exposed blocks were removed in the controlled-lighting luminescence-dating laboratory at the Desert Research Institute (DRI) in Reno, NV, USA. The collected samples are described in Supplementary Table S1.

Prior age estimates at Atapuerca

Gran Dolina

The presence of remains of the micromammal *Miomys savini* (a water vole) only below the upper 40–50 cm of TD8 constrains this and lower levels to ages older than ca. 500 ka (e.g., Carbonell et al., 1995, 1999; Laplana and Cuenca-Bescós, 1997). The first reliable numeric age estimate for the TD deposits was provided by Parés and Pérez-González (1995, 1999), who located the B–M boundary within upper TD7, thus constraining the hominin fossils

in the underlying TD6 to be older than 780 ka, and making them the oldest European hominin fossils (e.g., Carbonell et al., 1995; Bermúdez de Castro et al., 1997). Below unit TD7, and to the lowest unit TD1, only reversed paleomagnetic polarity was observed, with a suggestion (a single-sample data point) of an excursion or reversal in unit TD1. However, between TD7 and TD1, the paleomagnetic samples were collected at widely spaced intervals. If the excursion in TD1 represents the Jaramillo polarity-normal interval, then all deposits between TD7 and TD1 represent the time interval 780–990 ka (Hornig et al., 2002), including the hominin remains in TD6. Recently, the strata containing lithic artifacts at Elefante (site TE; Fig. 1) have been determined to be older than ca. 780 ka, possibly older than 1.07 Ma, based on paleomagnetic analyses (Parés et al., 2006).

Above horizon TD7, useful material for numeric dating has been difficult to obtain. Although the deposits at Gran Dolina (and Galería) contain carbonate that is in principle suitable for U-series dating, the open-system behavior of the isotopes has precluded stratigraphically systematic results (Bischoff, pers. comm.). We note, however, that two samples of vein fill within TD11 produced reasonable U-series age estimates of 240 ± 25 ka (average) (Bischoff, pers. comm.).

Combined ESR-U-series dating of ungulate tooth enamel from the Gran Dolina site (Falgüères et al., 1999, 2001) has produced the most systematic results so far for both Gran Dolina and Galería. Age estimates range from 300–400 ka in the upper part of TD10 to ca. 600 ka in TD8, and ca. 750 ka in TD6, below the B–M boundary (Fig. 2, left). Note that Falgüères et al. (1999, 2001) placed three of their samples well within unit TD11, but the correct stratigraphic placement apparently is near the top of the underlying TD10 (Fig. 8 in Pérez-González et al., 2001). Similarly, for the underlying TD10 samples of Falgüères et al. (1999, 2001), Pérez-González et al. (2001) provide a clearer stratigraphic location.

At nearby Sima de los Huesos (Fig. 1), U-series dating of carbonate (in situ speleothem) also was successful, producing a lower-limiting age estimate of ca. 350 ka, with a stratigraphically extrapolated probable age of 400–500 ka (Bischoff et al., 2003) for underlying human bones. That age estimate has been revised upward to a minimum of ca. 530 ka (Bischoff et al., 2007) by use of higher-precision U-series dating.

Galería

Very few dating results have been reported for Galería. The B–M boundary has been placed in the upper part of unit G1, 3.5 m above its base (Pérez-González et al., 1999, 2001). At the boundary between G1 and the overlying G2, a speleothem has yielded an ESR age of 317 ± 60 ka and a limiting U-series age estimate of greater than 350 ka (reviewed by Aguirre, 1994; Pérez-González et al., 2001). Within the lower part of the overlying G3 unit, speleothem has given an ESR age estimate of 256 ± 33 ka (Aguirre, 1994), while carbonate pieces and speleothem in the base of the overlying lower G5 unit have yielded ESR age estimates of 177 ± 23 ka and 211 ± 32 ka (summarized by Aguirre, 1994; Pérez-González et al., 1999, 2001; Falgüères et al., 2001) (Fig. 2, right).

Because the sediments at both Gran Dolina and Galería are older than ca. 100 ka and are directly non-eolian (not wind borne), they presumably would be difficult to date with luminescence sediment-dating methods. Nevertheless, we have applied luminescence methods to some of these deposits, motivated by the reported success of Berger (1994, 1995) and Berger et al. (1992) in attaining accurate TL ages from fine-silt-sized feldspars up to ca. 800 ka, and accurate IRSL ages up to ca. 400 ka (Berger, 2001) for several independently dated eolian fine-silt deposits (loess) from New Zealand and Alaska.

We reported early results in abstracts (Berger and Pérez-González, 2000; Berger et al., 2005). Since then, only a few other attempts have been reported to date karstic infill sediments by luminescence methods (Morwood et al., 2004; Turney et al., 2001; Prideaux et al., 2007a,b). Morwood et al. (2004) applied IRSL and red-TL procedures to cave infill sediments to obtain age estimates of ca. 14 ka and 35 ka stratigraphically bracketing a hominin skeleton. These results were obtained from slopewash sediments, but it was not reported whether the luminescence in the sediment grains was zeroed before deposition in the cave. Thus, the ca. 35 ka age estimate could greatly exceed the burial age. Stratigraphically comparable radiocarbon ages average ca. 18 ka. Turney et al. (2001) reported OSL (optically stimulated luminescence, or optical) ages in the range 40–50 ka for quartz grains from slopewashed sediments enclosing human artifacts. Here, the sediments originated mainly from eolian dunes outside the cave, thus ensuring adequate zeroing of the luminescence clock before grain burial in the cave. Similarly, Prideaux et al. (2007a) reported OSL ages for quartz from quartz-rich horizons in Cathedral Cave ranging from ca. 200 ka to ca. 500 ka. Although they did not comment on the likely origin of this quartz, they apparently presumed that the quartz was washed into the cave from originally eolian deposits. Prideaux et al. (2007b) reported OSL age estimates for cave-infill quartz grains ranging from ca. 30 ka to ca. 230 ka, but no mention was made of the origin of these grains. Because these grains were inwashed during flood-water episodes, it is unclear how well the luminescence clock might have been zeroed before grain burial. The single-aliquot results of Prideaux et al. (2007b) indicate poor clock-zeroing for most of the infill sediment.

Several reports of luminescence dating of other rock-shelter and cave-infill unheated sediments have been made, for which the cave bedrock is sandstone, micaceous schist, other noncarbonate rock, or calcarenite, and the deposits are younger than ca. 80 ka. In these other applications (Abeyratne et al., 1997; Feathers, 1997; Smith et al., 1997; Jacobs et al., 2003a,b), fine-sand-sized quartz has been employed, rather than the polymineral fine-silt fraction that we used. In most of these applications (Abeyratne et al., 1997; Feathers, 1997; Smith et al., 1997), the authors assumed their samples were entirely or dominantly eolian, and some conducted tests to check this assumption. Each group discussed the possibility that quartz from comminuted fragments of the cave walls and roof were incorporated into their samples. Feathers (1997) also allowed that local colluvium could have been admixed with eolian grains. These groups employed multi-aliquot procedures or early versions of many-grain single-aliquot procedures. On the other hand, Jacobs et al. (2003a,b) applied the more recent many-grain-single-aliquot and the single-grain procedures to sand-sized quartz grains that are clearly eolian and deposited within a calcarenite cave.

In all such applications, where the bedrock contains a significant fraction of quartz, and where the infill deposits are younger than 100–150 ka, the most accurate luminescence-dating approach is probably the single-grain-quartz approach, pioneered in such settings by Roberts et al. (1999) and advanced by Jacobs et al. (2003b). This approach enables, as can the few-grain-single-aliquot method (e.g., Prideaux et al., 2007b), statistical discrimination between eolian or last-zeroed quartz grains and older grains representing fragments of wall rock or colluvium containing significant relict luminescence signals at burial time. Equipment for the application of this modern approach was not available to us during this project, nor did we judge that the method would likely have been useful because of the expected antiquity of our cave sediments.

An age estimate equals paleodose divided by dose rate, which equals D_E/D_R , where D_E is derived from luminescence measurements. To derive D_E values, we used the total-bleach (TB) extrapolation-dose-response procedure for both TL and IRSL (Berger and Péwé, 2001; Berger, 2003), and also the “Australian slide” regenerative-dose-response procedure (Prescott et al., 1993) for both TL and IRSL. However, the ten samples analyzed with the slide procedure consistently yielded underestimates in D_E values compared to the TB extrapolation-dose-response procedure for most samples and its use was discontinued, with no results being reported here.

The TL and IRSL procedures we chose require adequate concentrations of K-feldspars in the fine-silt size fractions. Pérez-González et al. (1995, 1999) reported general estimates of feldspar concentrations within the cave deposits. To obtain localized estimates of feldspar concentrations and the ratios of compositional groups of feldspars—as well as quartz-feldspar ratios—from the identical ~0.5-mg aliquots used for luminescence dating, we conducted a few point-count SEM-EDX probe analyses.

To calculate dose rates, we used the laboratory procedures employed by Berger and Péwé (2001): commercial atomic absorption spectroscopy (AAS) for K_2O , and thick-source-alpha-particle-counting (TSAC; Huntley and Wintle, 1981) for U and Th count rates. Two supplementary procedures were used to check the AAS and TSAC results. In 1997, we employed a 2π -geometry, 4-channel portable, NaI(Tl)-crystal gamma spectrometer (e.g., Aitken, 1985) at some horizons. In 1998, several ~1-kg samples were collected specifically for intrinsic-Ge high-resolution gamma spectrometry (HRGS) at the University of Southern California (USC) (e.g., Luo et al., 2000).

Further details of the above methods pertaining to this project are provided in the online Supplementary material.

Results

The results of SEM-EDX analyses for mineral-concentration estimates are listed in Table 1. The images for sample ATP97-49 (Fig. 3) are representative of the selected samples. Most of the feldspar grains appear to be K-rich, with a small to negligible fraction in this example being Ca-rich (therefore dark or only slightly visible in Fig. 3). No albite (Na-rich feldspar) grains were identified. No zircon grains were detected. Zircon can produce problematic luminescence (e.g., Aitken, 1998). Overall, these results indicate that we have a usefully large fraction of K-feldspars

Table 1
Mineral percentages from SEM-EDX analyses of 4–11- μ m-diameter grains^a

Sample (ATP)	Total points	Quartz	K-Fs	Hybrid ^b Fs	Plagioclase	Lithics	Other ^c
<i>Galería</i>							
98-19	200	27.5	15.0	6.5	0	5.5	45.5
97-24	175	28.6	1.7	3.4	0	4.6	61.7
98-12	175	44.0	13.1	2.3	1.1	6.9	32.6
97-6	175	13.7	6.9	4.6	0	8.6	66.3
0.5							
<i>Gran Dolina</i>							
97-33	175	21.1	10.3	10.3	0	5.1	53.1
97-28	175	36.6	10.9	10.9	0	8.0	33.7
98-20	175	25.7	17.7	0	0	13.1	43.4
97-49	175	39.4	7.4	6.3	0	2.9	44.0
97-41	200	30.5	14.0	0	0	5.5	50.0

^a Samples were treated previously in HCl acid to dissolve carbonates and in H₂O₂ to destroy organic material.

^b This represents composite grains of feldspar (e.g., partially altered feldspars).

^c Generally, this category represents, in order: muscovite (or other Al-rich K-silicates) and minor-Fe clay minerals, with minor amounts of biotite and amphibole.

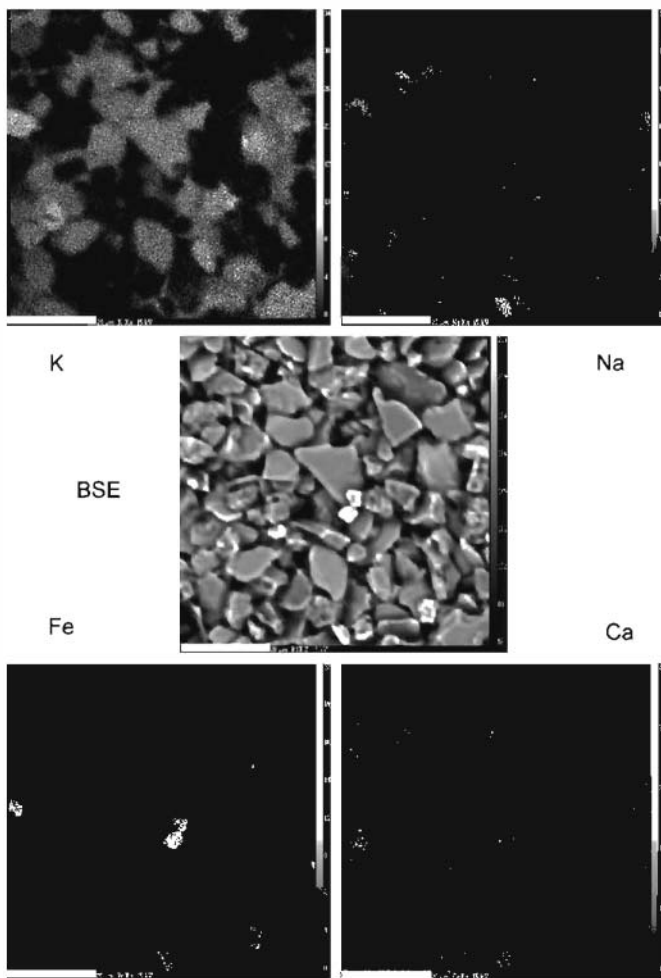


Fig. 3. Back-scatter-electron (BSE) SEM image (center) for siliciclastic grains of 4–11 μm diameters in an aliquot of sample ATP97-49 previously used for TL. Also shown are SEM-EDX element-map images of the same field of view tuned to detect K, Na, Fe, and Ca. Dark areas indicate grains having little or none of the target elements. White scale bars (lower left) are 20 μm .

(7–18%), little or no distinct plagioclase (<1%), and 21–44% quartz in this fine-silt size range. Because the TL from feldspars is 20–50 times brighter than that from quartz, the observed relative concentrations in Table 1 are favorable for TL dating. Moreover, as K-feldspars emit a generally stable signal at ~ 420 nm wavelengths (Berger, 1994; Krbetschek et al., 1997), we performed all TL experiments using this emission.

The K-, U-, and Th-concentration values calculated from in situ 4-channel gamma field spectrometry (“field γ ” hereafter) and from laboratory high-resolution gamma spectrometry (HRGS hereafter) are listed in Supplementary Table S2. These are compared to the laboratory values from AAS and TSAC, employed for all TD and TG samples. The results indicate general agreement among the methods. Moreover, from the HRGS data, there appears to be no significant (for dose-rate calculations) radioactive-decay-series disequilibria (e.g., Olley et al., 1996, 1997). For this reason and others outlined in the Supplementary materials, we chose to calculate the dose rates using only the AAS and TSAC data.

The age estimates for the several surface-soil samples (Table 2) are in the range 0.3–6.0 ka, indicating that, if such material washed into the past caves, it would have carried a relict age of less than 10 ka. Thus, even without additional (uncertain) clock zeroing during downslope transport, these results suggest that inwash sediment forming the matrix of the cave deposits carries

a statistically negligible systematic inherited-age offset for samples older than ca. 100 ka. Very likely, these low luminescence age estimates reflect the presence of modern and Holocene eolian grains (see online Supplementary material).

At Gran Dolina, two subadjacent samples of older terra rossa soil (samples 97-13; Supplementary Table S1 and Fig. 2) yielded a mean age of 55 ± 14 ka, consistent with their stratigraphic position at the top lip of the cave passage. The luminescence results from these samples are consistent with an eolian origin of silt grains within forming surface soils (incipient and developed terra rossa) and subsequent downslope migration into the caves.

Our judgment of the accuracy of the TL and IRSL age estimates for cave-matrix sediment depends upon comparisons both within the luminescence data sets and between the data sets, as well as comparisons with independent relative (e.g., stratigraphic) and numeric (e.g., U-series, ESR, and ESR-U-series) age criteria.

With IRSL data, we looked for constancy of plateau D_E values when we increased preheating temperatures above a certain level (e.g., above 140°C ; Ollerhead et al., 1994). An example of such constancy is shown in Fig. 4. We also looked for agreement between TL plateau D_E values and those from IRSL (e.g., Fig. 4). For several samples in Table 2, we calculated weighted (by inverse variance) mean and concomitant standard error of the mean (e.g., Topping, 1962) when such concordance in D_E values was observed. Particular anomalies in the TL D_E -temperature plots (e.g., above 350°C in Fig. 4) are discussed in the online Supplementary material.

Comparison between TL and IRSL D_E plots revealed that, at both Gran Dolina and Galería, IRSL D_E plateau values began to diverge from TL values for samples older than ca. 250 ka, with the IRSL values ceasing to increase with stratigraphic age. Figure 5 (inset) provides an example of this age underestimation from IRSL data. Consequently, for all samples older than ca. 250 ka, we consider that only the TL results provide reasonable age estimates. Furthermore, as discussed in the online Supplementary material, we selected TL D_E values only in the range below ca. 370°C in D_E -temperature plots. An example of this selection and the resultant age for the oldest sample analyzed is shown in Fig. 6 (inset). The imprecision of the resultant mean age arises from the relatively large uncertainties associated with the long dose-response extrapolations with such samples near the upper age limit of this dating procedure.

Based on the above reasoning, we plot in Fig. 2 only those age estimates in Table 2 denoted by two asterisks (**). In addition, for Galería, we noticed that some of the age estimates from within a given stratigraphic unit are statistically identical (at 67% confidence). We therefore plot only the weighted mean ages for those results in Fig. 2 (right).

Discussion

Stratigraphically anomalous TL ages

A likely explanation for the anomalous 1.34-Myr IRSL result in the middle of unit G4 (Fig. 2, right, sample 98-19) is discussed in the online Supplementary material. More relevant to our interpretation of overall results is the ca. 3-Myr TL result in unit G1, just above the B–M boundary. This sample (97-2) was intentionally collected from the interior facies. As such, an anomalous age overestimate was expected because the sediment grains were deposited and redeposited by subterranean streams and water flows, precluding any exposure to daylight since they entered the cave systems. This sediment entered the caves probably in the early Pleistocene or before (e.g., Carbonell et al., 1999; Parés and Pérez-González, 1999). Thus, when compared to the TL results from the overlying units G2 and G3, this TL result confirms the geological interpretation of the deposits as consisting of either clearly interior facies (unit G1) or clearly exterior facies (overlying units).

Table 2
Dose rate, IRSI, and TL data, with calculated ages for the Arapuerca samples

Stratigraphic unit	Sample ^a (ATP)	Water ^b (=C.05)	K ₂ O wt% (=C.05)	C ₂ ^c (ks ⁻¹ cm ⁻²)	C ₀ ^d (ks ⁻¹ cm ⁻²)	E value ^e (pGy cm ²)	Dose rate ^f (Gy/ka)	Preent ^h (%)	D ₀ ^g (Gy)	Age ⁱ (ka)
<i>Surface soils</i>										
0.5–3.0 cm	97-1					0.95 ± 0.09	3.0 ± 0.4 ^j	130/150	1.55 ± 0.15 (IR-120/2)	0.52 ± 0.08
0.5–3.0 cm	98-1					0.94 ± 0.11	3.0 ± 0.4 ^j	140	2.24 ± 0.21 (IR-420/1)	0.75 ± 0.11
0.5–3.5 cm	98-2					0.98 ± 0.08	3.0 ± 0.4 ^j	130/150	4.98 ± 0.23 (IR-420/2)	1.49 ± 0.22
0.5–3.0 cm	98-3					1.01 ± 0.21	3.0 ± 0.4 ^j	130/150	1.78 ± 0.15 (IR-420/2)	0.26 ± 0.04
10 cm	97-11					0.82 ± 0.07	3.0 ± 0.4 ^j	130/150	4.02 ± 0.14 (IR-120/2)	1.34 ± 0.20
40 cm	97-12					0.84 ± 0.05	3.0 ± 0.4 ^j	160	19.30 ± 0.65 (IR-420/2)	6.43 ± 0.97
<i>Galería</i>										
G4/mid	98-19	0.14	0.18	0.1319 ± 0.0032	0.0606 ± 0.0093	1.38 ± 0.26	0.606 ± 0.061	160/180	814 ± 64 (IR420/2)	1340 ± 170**
			0.11	0.0917 ± 0.0026	0.1334 ± 0.0051					
G4/base	97-14	0.25	2.04	0.660 ± 0.011	0.380 ± 0.037	0.82 ± 0.26	2.76 ± 0.27	180	1405 ± 240 (TL ₁)	510 ± 100
			1.27*	0.600 ± 0.058	0.347 ± 0.025			150/160	5.11 ± 5.1 (IR420/2)	185 ± 26**
G3/top	98-17	0.18	0.44	0.2287 ± 0.0041	0.123 ± 0.013	0.9 ± 0.2	1.460 ± 0.092	160	4.72 ± 44 (TL ₁)	
			1.99	0.4939 ± 0.0098	0.262 ± 0.034			160/180	1.77 ± 3.5 (IR420/2)	269 ± 27
									Weighted mean D₀ = 393 ± 28	269 ± 27
G3/mid	97-24	0.25	1.97	0.6940 ± 0.0084	0.381 ± 0.030	0.81 ± 0.45	2.62 ± 0.38	175	555 ± 71 (TL ₁)	224 ± 42
			0.93*	0.349 ± 0.070	0.184 ± 0.039			150/170	4.70 ± 8 (IR420/2)	
									Weighted mean upper G3 = 256 ± 23**	
G3/lower	98-15	0.18	0.55	0.3093 ± 0.0044	0.6285 ± 0.0095	1.06 ± 0.29	1.42 ± 0.12	175	580 ± 34 (TL ₁)	480 ± 48
			0.69	0.3307 ± 0.0062	0.177 ± 0.020			150/170/170	6.5 ± 2.5 (IR420/3)	
G3/base	97-19	0.18	1.92	0.816 ± 0.011	0.494 ± 0.041	1.48 ± 0.45	3.72 ± 0.42	180	1.030 ± 100 (TL ₁)	439 ± 66
			1.18*	0.492 ± 0.080	0.250 ± 0.032					
									Weighted mean base G3 = 466 ± 39**	
G2/mid	98-12	0.25	2.30	0.851 ± 0.012	0.364 ± 0.039	3.03 ± 0.73	3.04 ± 0.26	165	1.280 ± 150 (TL ₁)	422 ± 55**
			1.94*	0.576 ± 0.12	0.252 ± 0.061			175	3.03 ± 150 (IR550/1)	
G2/base	97-6	0.20	2.27	0.805 ± 0.012	0.430 ± 0.041	1.19 ± 0.49	3.92 ± 0.42	175/180	1.370 ± 310 (TL ₂)	503 ± 95**
			1.72	0.769 ± 0.011	0.399 ± 0.011			160	0.08 ± 3.5 (IR420/1)	
Above B–M	97-2	0.22	1.71	0.817 ± 0.013	0.480 ± 0.048	1.0 ± 0.2 ^k	3.50 ± 0.20	180	10.500 ± 6300 (TL ₁)	3000 ± 1800**
		0.19	1.79*	0.926 ± 0.010	0.476 ± 0.025					
<i>Gran Dolina</i>										
TD11/wedge	97-13a						3.0 ± 0.4 ^j	150	1.38 ± 32 (IR420/1)	46 ± 7
TD11/wedge	97-13b						3.0 ± 0.4 ^j	160	2.20 ± 3.1 (IR420/1)	76 ± 11
									Weighted mean wedge = 55 ± 14**	
TD11/upper	97-37	0.15	1.32	0.4407 ± 0.0080	0.271 ± 0.029	1.22 ± 0.17	2.30 ± 0.16	170	3.5 ± 1.34 (IR550/1)	198 ± 19**
								180	4.73 ± 74 (IR420/1)	240 ± 44**
								170	4.87 ± 4.7 (TL420/1)	
									Weighted mean "420" D₀ = 484 ± 35	
TD11/mid	97-33	0.20	1.73	0.6057 ± 0.0093	0.311 ± 0.032	1.74 ± 0.75	3.36 ± 0.45	160/175	810 ± 100 (IR420/2)	
			1.56*	0.5153 ± 0.0060	0.305 ± 0.010					
TD10/upper	97-45	0.20	1.93	0.965 ± 0.014	0.420 ± 0.046	1.97 ± 0.85	5.24 ± 0.79	180	1090 ± 120 (TL ₁)	208 ± 39
			2.44	1.015 ± 0.12	0.460 ± 0.030					
TD10/upper	97-28	0.20	2.33	0.901 ± 0.011	0.378 ± 0.037	0.72 ± 0.13	3.39 ± 0.20	175/185	944 ± 91 (TL ₂)	278 ± 31
			1.60*	0.771 ± 0.010	0.339 ± 0.020					
			0.08*	0.0866 ± 0.0042	0.034 ± 0.010					
TD10/upper	98-23	0.20	2.33	0.901 ± 0.011	0.378 ± 0.037	0.78 ± 0.30	3.51 ± 0.31	195	670 ± 240 (TL ₁)	191 ± 70
			1.60*	0.771 ± 0.010	0.339 ± 0.020					
			0.08*	0.0866 ± 0.0042	0.034 ± 0.010					
									Weighted mean upper TD10 = 244 ± 26**	
TD10/mid	98-20	0.35	2.26	0.918 ± 0.011	0.540 ± 0.042	0.54 ± 0.14	2.82 ± 0.25	175/185	1.212 ± 1.25 (TL ₂)	430 ± 59**
			1.22*	0.58 ± 0.25	0.28 ± 0.13			160	5.80 ± 80 (IR420/1)	
								165	3.68 ± 9.5 (IR550/1)	
TD9	97-53	0.20	1.91	0.821 ± 0.12	0.459 ± 0.043	1.32 ± 0.24	3.97 ± 0.37	190	19.20 ± 490 (TL ₁)	480 ± 130**
		0.15	0.88*	1.18 ± 0.28	0.154 ± 0.020			190	7.50 ± 140 (IR420/1)	
								185	7.10 ± 2.50 (IR550/1)	

TD8/base	97-49	0.31	1.99	0.944 ± 0.012	0.409 ± 0.042	1.0 ± 0.2 ^k	3.21 ± 0.23	185	2630 ± 410/TL/1	816 ± 140 ^{**}
		0.18	1.02	0.533 ± 0.035	0.178 ± 0.028					
TD7/base	97-41	0.21	1.47	0.4962 ± 0.0090	0.281 ± 0.033	1.0 ± 0.2 ^k	2.48 ± 0.19	195	2280 ± 270/TL/1	
		0.26	1.54 [*]	0.66 ± 0.11	0.294 ± 0.016			200	2840 ± 560/TL/1	
Weighted mean D_e = 2390 ± 240										

Note: For some samples, IRSL ages are not listed in the age column (simply divide D_e by dose rate) because the D_e values are considered to be inaccurately too low, as discussed in the text and the online Supplementary material. The laboratory sample name. All samples are in stratigraphic order. All luminescence analyses are on the 4–11- μ m-diameter (fine silt) feldspar-quartz-mixture grains. The second row in columns 3–6 for some samples provides data to calculate the gamma dose rate from sediment within a distance 20–30 cm above and below the luminescence sample. Where absent, the data in these columns of the second row are statistically indistinguishable from those in the first row. The third row for sample 97-28 denotes analyses for limestone rock that is representative of ~30% of the volume of material contributing to the γ dose rate around this sample. The geometric volume contribution from such rock is incorporated into the dose-rate calculation.

^b Estimated past average ratio of weight of water/weight of dry sample. This estimate is 50–60% of the measured saturation value. Uncertainties here and elsewhere are reported at the 67% confidence level. Samples: 97-24, 98-12, and 97-3 have respective uncertainties of ±0.10, 0.08, and 0.03.

^c For the samples denoted with an asterisk (*), the uncertainties are respectively: ±0.40, 0.25, 0.25, 0.11, 0.08, 0.07, 0.08, 0.07, 0.06, 0.03, 0.22, 0.14.

^d Total and thorium count rates from finely powdered samples for thick-source-alpha-particle-counting (TSAC) method (Huntley and Wintle, 1981). C_T = C_α - C_{Th}. These values are inserted directly into the age equations of Berger (1988).

^e Alpha effectiveness factor (Huntley et al., 1988; Berger, 1988).

^f Calculated with the conversion factors given by Adamiec and Aitken (1998), using equations of Berger (1988), and includes a small cosmic-ray component estimated from the data of Prescott and Hutton (1988). These cosmic-ray components are (respectively down the table from the Galería subheader): 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02, 0.08 ± 0.02.

^g Each replicate (see Footnote h) D_e experiment had a preheating of two-day duration.

^h IRSL was detected at the 420 ± 20 nm spectral region (bandpass 390–470 nm at 1% cut) and 550 ± 30 nm (bandpass 510–600 nm at 1% cut) using optical glass filters specified in Berger and Péwé (2001). The number of replicate "total-bleach" experiments used to obtain the stated D_e value is given after the last slash.

ⁱ D_e divided by dose rate. Values marked with two asterisks (**) are age estimates plotted in Fig. 2. In addition, for Gran Dolina (excepting 55 ± 14 ka), individual TD10/upper TL age estimates are plotted in Fig. 7.

^j Because other dose-rate parameters were not measured for these samples (they were not as important as estimates of the D_e values for these surface samples and the top-most cavern-filling Dolina samples), these dose rates are estimated from the range of other samples to provide rough estimates of apparent ages.

^k These are estimates, based on measured values for other samples in this table because the α dose-response was saturated for these samples.

As discussed above, there have been very few independent numeric ages for these cave deposits. The ESR and U-series results mentioned above are summarized in Fig. 2 (right). The luminescence age estimates are consistent with those results and fill in stratigraphic age gaps. Moreover, the luminescence age estimates suggest the existence of a previously unsuspected time-stratigraphic gap within unit G3 in the interval ca. 250–400 ka. Additional, more precise (stratigraphically and analytically) dating is needed to confirm this suggestion.

Nonetheless, these results at this site now suggest that unit G3 and at least the upper part of unit G2 are likely younger than the rich hominin-fossil-bearing deposits in nearby Sima de los Huesos (Fig. 1), recently estimated to be at least 530 ka in age (Bischoff et al., 2007). The imprecise result (503 ± 95 ka) for the base of unit G2 adds little to this between-site comparison, and the entire unit G2 may be only ca. 450 kyr old (the weighted mean of the two results in Fig. 2, right, is 442 ± 48 ka). Nevertheless, this new relative chronostratigraphy for the two sites may help in the interpretation of tool development in the middle Pleistocene because Galería is rich in fauna and artifacts.

Age comparisons at Gran Dolina

The stratigraphically correct locations (Pérez-González et al., 2001) of the combined ESR-U-series age estimates from ungulate teeth (Falguères et al., 1999, 2001) are plotted in Fig. 2 (left). The luminescence ages above unit TD10 appear reasonable. Interestingly, two unpublished U-series age estimates for vein fill in TD11 of mean 240 ± 25 ka (Bischoff, pers. comm.) are consistent with the luminescence results from TD11.

To provide a clearer graphical comparison of the various results from Gran Dolina, we plot in Fig. 7 the mean ESR-U-series age estimates from Fig. 2 with the main luminescence results from Fig. 2, as well as the individual TL ages from upper TD10. The TL results in the central Fig. 7 may be interpreted in two ways: (1) the TL results from upper TD10 and the oldest ESR-U-series results (in TD6) are inaccurate (too young); or (2) two of the ESR-U-series mean ages are inaccurate (the one at ~16 m in uppermost TD10, and the one below the Aurora stratum). The second interpretation implies that the luminescence ages from upper TD10 are accurate and that the apparent 250–400-kyr stratigraphic age gap observed at Galería occurs also at Gran Dolina. Apart from the possibility that the TL results from upper TD10 are too young because of unrecovered anomalous-fading components (see discussion in online Supplementary material), further studies are required to determine which of these two interpretations for upper TD10 is correct.

The TL ages from units below upper TD10 are consistent with both the ESR-U-series age estimates and the presence of the B–M boundary in upper TD7. This consistency gives us confidence in the TL results. The two oldest TL results (lower unit TD8 and unit TD7) and the ESR-U-series ages from the middle of TD8 are also consistent with the biostratigraphic age-control (mentioned above) in upper unit TD8 that is provided by the last occurrence of remains of the vole *M. savini*. This last-presence datum assigns lower deposits to ages older than ca. 500 ka. The data in Figs. 2 (left) and 7 suggest that the ESR-U-series ages (mean 730 ± 63 ka) from unit TD6 are underestimations.

In summary, the TL results from the units TD8 and TD9 above the B–M boundary at Gran Dolina support the ESR-U-series results from TD8 and establish clearly that unit TD8 represents a time interval corresponding to the lower-limiting age (ca. 530 ka; Bischoff et al., 2007) of the several remarkable hominin fossils in Sima de los Huesos.

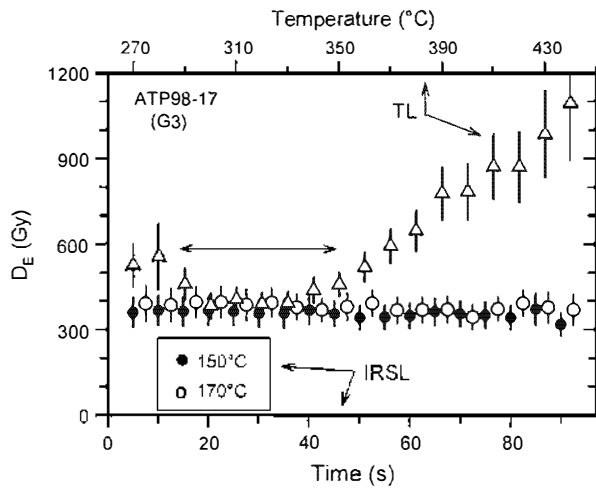


Fig. 4. D_E - t plots for IRSL and a D_E - T plot for TL from sample ATP98-17 from unit G3. The horizontal arrow denotes the temperature range used for calculating a mean TL D_E value, and thus a TL age for this sample.

Implication of the oldest TL result at Gran Dolina

The age estimate of ca. 960 ka for sample 97-41 (Figs. 2 and 7), although imprecise, may be the oldest TL feldspar result in the world that is consistent with stratigraphic and other evidence (in this case, paleomagnetic). Taken at face value, this result implies that the hominin fossils in the TD6 unit at Gran Dolina are probably about 900 kyr old or somewhat older (Fig. 7).

From a strictly numerical-dating viewpoint, it is logical to extrapolate the main age-depth trend in Fig. 7 and to consider the implications. Linear extrapolation of the main trend is approximated by line *a*, which is constrained to pass through the intersection of the main trend with the Aurora stratum to the age-depth point of 1.8 Ma at 0.9 m. At 0.9 m, a single-sample normal-polarity data point is reported (Parés and Pérez-González, 1999) to be either an excursion or a normal-polarity interval within the longer Matuyama reversal interval. The Olduvai normal-polarity interval occurs at ca. 1.8 Ma (Flörng et al., 2002). If the normal-polarity datum at 0.9 m represents the Jaramillo interval (as preferred by Parés and Pérez-González, 1999), then trend *b* is more probable.

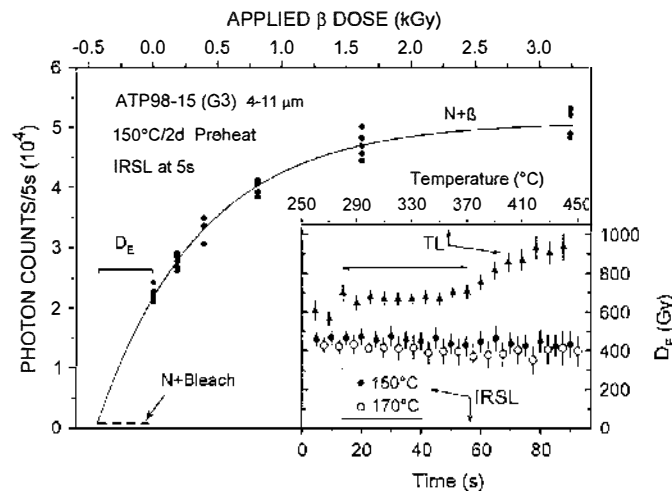


Fig. 5. Total-bleach dose-response D_E - t and D_E - T plots for sample ATP98-15 (unit G3). Here, the TL data are fitted to a saturating-exponential (E) regression model.

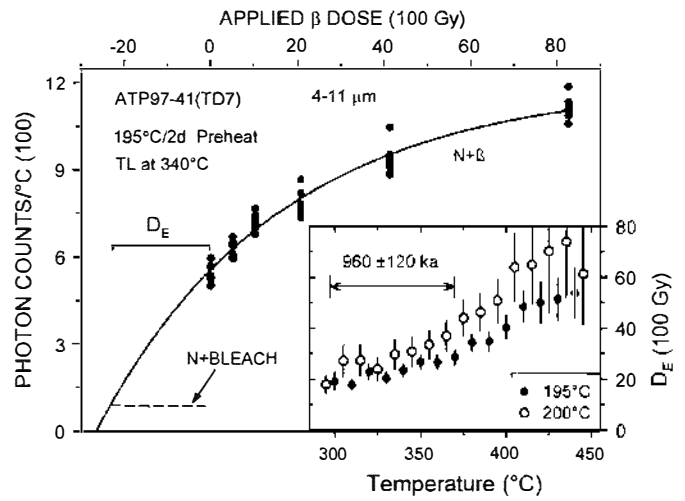


Fig. 6. Total-bleach dose-response and D_E - T plots for the oldest luminescence sample at Gran Dolina. Here, eight aliquots were used at each dose point, with an E regression model.

Stratigraphic age-depth trend *a* implies that sediments at Gran Dolina are preserved at least since ca. 1.8 Ma. How realistic is this? At the nearby Elefante site (Fig. 1), the greater portion of the ~25 m of exposed sediments was deposited before 0.8 Ma, likely in the interval 0.8–1.8 Ma, based on paleomagnetic reversal stratigraphy and biostratigraphic evidence (Parés et al., 2006). In the latter case, the recovery of Mode I (Oldowan) lithic-tool fragments in sedimentary units TE9 to TE13 (numbered bottom to top; Parés et al., 2006) and a hominin premolar and mandible in unit TE9 (Carbonell et al., 2008) clearly establishes that humans were present in this area before at least 1.1 Ma. Based on the insectivore assemblage in units TE8 to TE14, Cuenca-Bescós and Rofes (2004) considered this sequence to be older than ca. 1.2 Ma, while Rofes and Cuenca-Bescós (2006: 313) deduced that the whole faunal assemblage in this sequence “indicates a warm period that could be correlated with the north European Waalian” period (ca. 1.25–

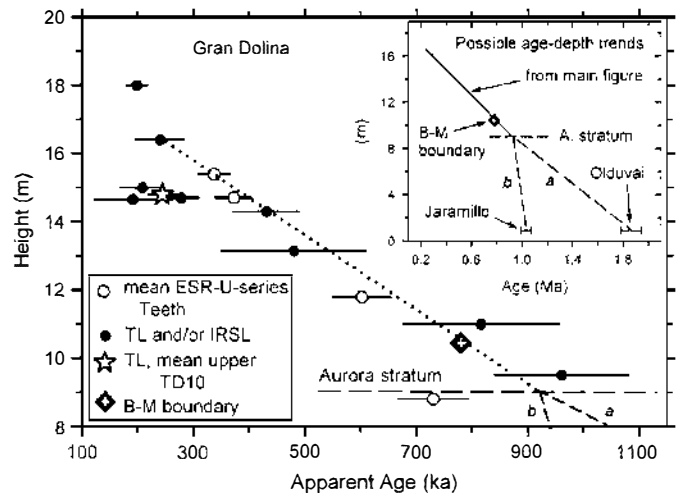


Fig. 7. In the main figure, a plot of TL and IRSL ages from Table 2 denoted with two asterisks (**) (see Footnote i in Table 2) and mean ESR-U-series ages from ungulate teeth at Gran Dolina, using the stratigraphic locations for the ESR-U-series results from Pérez-González et al. (2001). The hominin fossils recovered here occur in the Aurora stratum. The dotted line is a conjectured age-depth trend. The inset figure shows this trend line, the Aurora stratum, the chronostratigraphic location of the B-M reversal boundary, and two postulated sedimentation-rate trends (*a* and *b*) discussed in the text.

1.50 Ma). Another datum relevant to discussion of trend *a* is the observation (García and Arsuaga, 1999) of remains of the hyaenid *Crocota crocuta* in units TD3–TD4 (~4 m in Fig. 7, inset) at Gran Dolina. Trend *a* (Fig. 7) implies that the remains of *C. crocuta* in TD3–TD4 represent a first appearance in western Europe at ca. 1.4–1.5 Ma. This is not unreasonable given that *Crocota* was already present on the Arabian Peninsula by ca. 1.5 Ma (García and Arsuaga, 1999). Finally, trend *a* removes the need for an explanation for the apparently sudden change in average sedimentation rate around the level of the B–M boundary that is required by trend *b*.

Set against this trend-*a* extrapolation are the arguments in support of trend *b*. Trend *b* is consistent with the interpretation of Parés and Pérez-González (1999) and Pérez-González et al. (2001) that the single-sample paleomagnetic datum at 0.9 m in Gran Dolina represents either the Jaramillo normal-polarity interval (ca. 1.1 Ma) or the slightly older (ca. 1.18 Ma; Horng et al., 2002) Cobb Mountain normal-polarity interval.

Trend *b* is supported by previous biochronological interpretations, though these necessarily have low time resolution. The biochronological interpretations of carnivores and small mammals, among others (Cuenca-Bescós et al., 1999; García and Arsuaga, 1999, 2001; Antoñanzas and Cuenca-Bescós, 2002; Cuenca-Bescós and García, 2007), indicate that the Gran Dolina strata near and below TD6 (which contains hominin remains) have a faunal assemblage “a bit more advanced than the Jaramillo faunas in Vallonnet and Les Valerots (France) or Untermassfeld (Germany)” (G. Cuenca-Bescós, pers. comm.), and thus somewhat younger than Jaramillo age. Moreover, the TD6 and stratigraphically nearby faunal assemblages appear to be younger than those at the Elefante site, which we stated above to be of probable pre-Jaramillo age.

Moreover, trend *b* is more consistent with indirect geomorphic evidence than is trend *a*. Trend *a* implies that unit TD1 is of Olduvai age, but this is apparently contradicted by geomorphic evidence. Because TD1 deposits are of vadose origin (above base level), and their elevation (~990 m above sea level) correlates to an elevation just above the +60–67-m (above the present alluvial plain) Pleistocene fluvial terraces of the nearby Arlanzón River (e.g., Parés and Pérez-González, 1999), then trend *a* implies that these terraces at +60–67-m elevation also must be of Olduvai age. However, Benito-Calvo et al. (2008) interpreted these +60–67-m terraces as of Brunhes-Matuyama-reversal age (~0.8 Ma), much younger than the Olduvai. The paleomagnetic-polarity directions from these +60–67-m terraces are reversed, while the next younger (lower, +50–54 m) and older (higher, +70–78 m) terraces contain normal-polarity directions (Benito-Calvo et al., 2008). Other terraces in the region under +60–67 m in elevation have a middle Pleistocene age based on the fauna (Sesé et al., 2000) and lithics (Santonja and Pérez-González, 2002) of middle-Pleistocene-age associations. In summary, this indirect geomorphic evidence suggests that level TD1 at Gran Dolina is significantly younger than Olduvai age. Efforts have begun by the Atapuerca research team to date the regional terraces (e.g., by cosmogenic nuclides), and this effort should help to settle the question of which extrapolated age-depth trend is more realistic.

Additional evidence supporting either of trend hypothesis *a* or *b* could logically come from more detailed dating studies within the lower units at Gran Dolina. Several of the paleomagnetic samples below the Aurora stratum analyzed by Parés and Pérez-González (1999) were spaced 1.0–1.5 m apart. Thus, the presence of Jaramillo-age sediments somewhat below TD6 could have been undetected. Improved precision of luminescence dating within TD7 and older (if possible) exterior-facies strata at Gran Dolina also would be useful. Finally, application of cosmogenic-nuclide dating of exterior-facies cave-sediment quartz grains (e.g., Anthony and Granger, 2007; Carbonell et al., 2008) has promise for these lower strata.

Conclusions

Luminescence dating is applicable to karstic infill sediments if there is a plausibly eolian origin to inwashed silt grains. With the dose rates measured at Atapuerca, multi-aliquot TL dating of feldspar-bearing fine silt is presently more reliable for samples older than ca. 200 ka than is multi-aliquot IRSL dating. Furthermore, comparisons of our TL results with stratigraphic constraints (viz., the B–M geomagnetic-reversal boundary at 780 ka), independent stratigraphic microfossil evidence at Gran Dolina, and ESR-U-series dating suggest that our preferred TL ages are generally accurate, and that reliable TL dates up to ca. 1 Ma can be obtained from fine-silt feldspars in some areas. Because of lack of access to the appropriate luminescence-reading instrument systems during this project, we have not investigated the usefulness of single-aliquot quartz grains in dating samples at Atapuerca, but we are skeptical of its potential for most of the strata (those older than ca. 200 ka). The sediments at Atapuerca have dose rates 5–7 times higher than those for the quartz-rich sediments analyzed by Rhodes et al. (2006) and Pridoux et al. (2007a), who reported quartz ages up to 500–900 ka.

The TL age-depth seriation at Gran Dolina suggests that, as at Galeria, the sediments and associated remains representing the time interval 250–400 ka are missing, but at Gran Dolina, the critical luminescence results in upper TD10 could be too young, and this apparent gap may be only an artifact there.

The TL dating results from Galeria suggest that faunal and lithic remains in stratigraphic units lower G3 and G2 at ca. 450 ka are younger than the remarkable hominin remains from the nearby Sima de los Huesos site (>530 ka). Our oldest TL result (960 ± 120 ka) for unit TD7 at Gran Dolina implies that the probable age of the hominin remains in the underlying TD6 horizon lies in the age range 900–950 ka. Since much of the TD6 unit represents an interglacial paleoclimatic stage (e.g., Sánchez-Marco, 1999), “perhaps very similar to that of the modern Cantabrian habitat” (Cuenca-Bescós et al., 2005: 285), our results (Fig. 7) imply an association of the hominin remains with Marine Isotope Stage (MIS) 25, a relatively warm and humid interglacial period. Previously, on the basis of an overinterpretation of the available ESR-U-series age estimates (e.g., Fig. 2, left, and Fig. 7) and biostratigraphic data, Cuenca-Bescós et al. (2005) suggested a correlation of the Aurora stratum with MIS 21.

The present results have important implications for reconstructing the history of human evolution in Europe. These results suggest an increase in the temporal gap between the TD6 hominins (*Homo antecessor*) and the middle Pleistocene hominins (*Homo heidelbergensis*). Cuenca-Bescós et al. (2005: 285) concluded that the presence of *H. antecessor* in TD6 “may reflect the ecological preferences of this species to the warmer and more equable conditions of southern Europe” whereas “the general opening of the landscape ... characterized in Atapuerca by the small-mammal association of TD10, could have favored the dispersal” of later hominins (e.g., *H. heidelbergensis*) across Europe. Our results also suggest the need for future studies of the possible phylogenetic relationship between these species (Bermúdez de Castro et al., 1997, 2003; Martín-Torres et al., 2007) and of the possible evolutionary events that occurred in Europe prior to 400 ka. Furthermore, the possible temporal coincidence between the TD6 hominins and the Ceprano calvaria (Ascenzi et al., 1996) offers another interesting debate about their possible relationship, especially if the species *Homo cepranensis* (Mallegni et al., 2003) is also accepted.

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Appendix A. Supplementary data

Supplementary material associated with this article can be found in the online version, at doi:10.1016/j.jhevol.2008.02.012.

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